

# Spectral degrees of cross-polarization of stochastic anisotropic electromagnetic beams in modified non-Kolmogorov atmospheric turbulence

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Based on the extended Huygens–Fresnel principle, variations in generalized Stokes parameters of stochastic anisotropic electromagnetic beams propagating through modified non-Kolmogorov atmospheric turbulence have been analyzed. The changes in generalized Stokes parameters with different turbulence parameters and source parameters were analyzed first. After that, the distributions of the spectral degrees of cross-polarization (SDCP) of isotropic beams and anisotropic beams were simulated. The results show that the profiles of distribution of SDCP of these two kinds of beams are very different in the near field, and will fluctuate through the propagation in atmospheric turbulence, but at last, when the propagation distance is long enough, the difference in the source makes a slight difference in the final profiles of SDCP in the output plane. They mainly depended on the turbulence perturbation, and in the weak turbulence, the profiles of final distribution show more flatter features.

Keywords: quasi-homogenous source, modified non-Kolmogorov turbulence, generalized Stokes parameters.

## 1. Introduction

The concept of cross-polarization of stochastic electromagnetic beam-like fields was first introduced by ELLIS and DOGARIU where the parameters are called the mutual degree of cross-polarization (MDCP) [1] and the quantity introduced in Volkov's paper is called the spectral degree of cross-polarization (SDCP) [2] which represent the correlation between the intensity fluctuations in terms of the degree of coherence and of the other two-point correlation properties of the beam. Recently, LI and PU have studied the degrees of cross-polarization of electromagnetic beams propagation in turbulence media and uniaxial crystals [3, 4]. PU has found that the behavior of this quantity, not

being similar to the behavior of the spectral degree of coherence or the ordinary degree of polarization, has been affected by all the parameters of the source. After the concept of a partially coherent Gaussian pulse was introduced [5], the spectrally partially coherent properties of such pulsed beams and the propagation of SDCP of stochastic electromagnetic pulsed beams in free space have been studied by DING and LAJUNEN [6–8]. Very recently, the statistic properties of stochastic anisotropic electromagnetic beams propagation in the atmosphere and oceanic water have been studied by DU and ZHOU [9, 10]. They have found that the changes in the degree of polarization of anisotropic beams are different from those of isotropic beams, because they have no self-reconstructed properties. On the other hand, atmospheric turbulence is one of the most important turbulences which has a significantly degrading impact on the quality of imaging and laser communication systems. The classical Kolmogorov turbulence spectral model failed to predict the result of recent experiments in certain portions of the atmosphere. The propagation of light beams through the turbulence in portions of the troposphere and stratosphere may obey non-Kolmogorov statistics [11]. XUE *et al.* generalized the modified turbulence spectral model in non-Kolmogorov atmosphere turbulence which considers the finite turbulence in inner and outer scales and has a general spectral power law value in the range of 3 to 5, instead of the standard power law value of 11/3 [12]. CUI *et al.* obtained the long exposure turbulence modulation transfer function (MTF) based on this model [13].

To our best knowledge, the influence of turbulence perturbation on the evolution of SDCP of stochastic beams is not clearly pointed out in the articles we mentioned above, there are a few reports about the propagation properties of the electromagnetic beams traveling through modified non-Kolmogorov turbulence. In this paper, we use generalized Stokes parameters of anisotropic stochastic electromagnetic beams to calculate the SDCP of anisotropic beams propagation through modified non-Kolmogorov turbulence. The results are illustrated by numerical examples. Basing on these results, we have discussed the changes in SDCP distributions under different turbulence conditions and source parameters. The behaviors of SDCP of such beams propagating in turbulence media were summarized at last.

## 2. Anisotropic electromagnetic beams and modified non-Kolmogorov turbulence model

The elements of the cross-spectral density matrix (CSDM) for a random, statistically stationary electromagnetic beam generated by the Gaussian source are given by:

$$W_{ij}^{(0)}(\boldsymbol{\rho}'_1, \boldsymbol{\rho}'_2, \omega) = A_i A_j B_{ij} \exp\left\{-\frac{\boldsymbol{\rho}'_1{}^2}{4\sigma_i^2} - \frac{\boldsymbol{\rho}'_2{}^2}{4\sigma_j^2}\right\} \exp\left\{-\frac{|\boldsymbol{\rho}'_2 - \boldsymbol{\rho}'_1|^2}{2\delta_{ij}^2}\right\} \quad (1)$$

where the coefficients  $A_i$ ,  $A_j$ ,  $B_{ij}$  and variances  $\sigma_i$ ,  $\sigma_j$  and  $\delta_{ij}$  are both independent of the position but dependent on the frequency;  $\boldsymbol{\rho}'_1$  and  $\boldsymbol{\rho}'_2$  are two points on the source plane. The degree of polarization is the same at every point across the source plane if

$\sigma_i = \sigma_j$ , but if  $\sigma_i \neq \sigma_j$ , the source is not uniformly polarized. For convenience, Eq. (1) can be rewritten as:

$$W_{ij}^{(0)}(\boldsymbol{\rho}'_{12}, \omega) = A_i A_j B_{ij} \exp \left\{ -\frac{ik}{2} \boldsymbol{\rho}'_{12}{}^T \mathbf{M}'_{ij}{}^{-1} \boldsymbol{\rho}'_{12} \right\} \quad (2)$$

where  $\boldsymbol{\rho}'_{12}{}^T = (\boldsymbol{\rho}'_1{}^T, \boldsymbol{\rho}'_2{}^T) = (x'_1, y'_1, x'_2, y'_2)$ , and

$$\mathbf{M}'_{ij}{}^{-1} = \begin{bmatrix} A & 0 & C & 0 \\ 0 & A & 0 & C \\ C & 0 & B & 0 \\ 0 & C & 0 & B \end{bmatrix}$$

while  $A = -\frac{i}{2k} \sigma_i^{-2} - \frac{i}{k} \delta_{ij}^{-2}$ ,  $B = -\frac{i}{2k} \sigma_j^{-2} - \frac{i}{k} \delta_{ij}^{-2}$ ,  $C = \frac{i}{k} \delta_{ij}^{-2}$ .

The CSDM of stochastic beams propagation through turbulence can be derived by the extended Huygens–Fresnel principle:

$$W_{ij}(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2, z; \omega) = \frac{k^2}{4\pi^2 z^2} \iint d^2 \boldsymbol{\rho}'_2 \iint d^2 \boldsymbol{\rho}'_1 W_{ij}^{(0)}(\boldsymbol{\rho}'_1, \boldsymbol{\rho}'_2, \omega) \times \\ \times \exp \left\{ -ik \frac{(\boldsymbol{\rho}_1 - \boldsymbol{\rho}'_1)^2 - (\boldsymbol{\rho}_2 - \boldsymbol{\rho}'_2)^2}{2z} \right\} \langle \exp \{ \psi^*(\boldsymbol{\rho}_1, \boldsymbol{\rho}'_1) + \psi(\boldsymbol{\rho}_2, \boldsymbol{\rho}'_2) \} \rangle_m \quad (3)$$

where  $\boldsymbol{\rho}_1, \boldsymbol{\rho}_2$  are two observation points on the output plane,  $z$  is the propagation distance and  $k = 2\pi/\lambda$  is the wave number of the beam. The last term describes the correlation function of the complex phase perturbed by random medium, the subscript  $m$  denotes the average over medium realization, and it takes the form:

$$\langle \exp \{ \psi^*(\boldsymbol{\rho}_1, \boldsymbol{\rho}'_1) + \psi(\boldsymbol{\rho}_2, \boldsymbol{\rho}'_2) \} \rangle = \\ = \exp \left\{ -4\pi^2 k^2 z \int_0^\infty \kappa \Phi_n(\kappa) \left[ 1 - J_0 \left| (1 - \xi)(\boldsymbol{\rho}'_1 - \boldsymbol{\rho}'_2) + \xi(\boldsymbol{\rho}_1 - \boldsymbol{\rho}_2) \right| \kappa \right] d\kappa d\xi \right\} \quad (4)$$

where  $J_0$  stands for the 0th order Bessel function. Under the approximation which was used in [14], Eq. (4) is simplified to:

$$\langle \exp \{ \psi^*(\boldsymbol{\rho}_1, \boldsymbol{\rho}'_1) + \psi(\boldsymbol{\rho}_2, \boldsymbol{\rho}'_2) \} \rangle = \\ = \exp \left\{ -\frac{\pi^2 k^2 z T}{3} \left[ (\boldsymbol{\rho}'_1 - \boldsymbol{\rho}'_2)^2 + (\boldsymbol{\rho}'_1 - \boldsymbol{\rho}'_2)(\boldsymbol{\rho}_1 - \boldsymbol{\rho}_2) + (\boldsymbol{\rho}_1 - \boldsymbol{\rho}_2)^2 \right] d\kappa \right\} \quad (5)$$

where  $T = \int_0^\infty \kappa^3 \Phi_n(\kappa) d\kappa$  denotes the intensity of the turbulence atmosphere which can only be determined by the spatial power spectrum of the refractive index fluctuations.

In the last decade, several researchers studied the propagation properties of beams traveling through non-Kolmogorov turbulence. Toselli has researched the long-term beam spread, scintillation index, probability of fail and mean bit error rate as a variation of the spectrum exponent of non-Kolmogorov turbulence [15]. GERÇEKCIOLU and BAYKAL scrutinized the behavior of intensity fluctuations of a flatted-topped beam in non-Kolmogorov weak turbulence [16]. CUI *et al.* investigated the influence of power law on the irradiance scintillation of a Gaussian-beam wave [17]. KOROTOKOVA and SHCHEPAKINA revealed the dependence of spectral shifts and switches in optical stochastic beams propagating through non-Kolmogorov media on the slope of the power spectrum of fluctuations in the refractive index [18]. The temporal power spectrum of the irradiance and the AOA fluctuations of plane and spherical waves in non-Kolmogorov turbulence whose power law varies from 3 to 4 were derived by DU *et al.* [19].

The general modified atmospheric spectrum is a turbulence spectral model which considers the influence of finite, inner and outer scales. For non-Kolmogorov turbulence, it has the form:

$$\Phi_n(\kappa, \alpha) = A(\alpha) \tilde{C}_n^2 \kappa^{-\alpha} f(\kappa, l_0, L_0, \alpha) \tag{6a}$$

$$f(\kappa, l_0, L_0, \alpha) = \left[ 1 - \exp\left(-\frac{\kappa^2}{\kappa_0^2}\right) \right] \left[ 1 + a_1 \left(\frac{\kappa}{\kappa_l}\right) - b_1 \left(\frac{\kappa}{\kappa_l}\right)^{7/6} \right] \exp\left(-\frac{\kappa^2}{\kappa_l^2}\right) \tag{6b}$$

where  $\kappa_0 = C_0/L_0$ ,  $\kappa_l = c(\alpha)/l_0$ ,  $c(\alpha)$  is the scaling constant,  $L_0, l_0$  are the out scale and the inner scale of turbulence. The intensity of modified non-Kolmogorov turbulence can be obtained by the numerical integral method of Matlab 2011a. Figure 1 shows a typical distribution of the intensity of the turbulence with power law. At first,  $T$  increases when the power law increases, and gets the maximum value at  $\alpha = 3.11$ , then it will decrease when the power law increases. For the homogeneous and isotropic turbulence, the power law  $\alpha$  varies from 3 to 4 when Markov's approximation is used.

From Equations (1)–(6), the cross-spectral density matrix of the output plane of the beam propagating through turbulence can be obtained by a tedious integral, what has been done in the previous work [5, 6]

$$W_{ij}(\mathbf{p}_{12}, z, \omega) = A_i A_j B_{ij} \left[ \text{Det}(\bar{\mathbf{I}} + \bar{\mathbf{B}}\bar{\mathbf{P}} + \bar{\mathbf{B}}\mathbf{M}'_{ij}{}^{-1}) \right]^{-1/2} \times \exp\left\{ \frac{ik}{2} \mathbf{p}_{12}^T \left[ (\bar{\mathbf{B}}^{-1} + \bar{\mathbf{P}}) - (\bar{\mathbf{B}}^{-1} - \frac{1}{2}\bar{\mathbf{P}})^T (\bar{\mathbf{B}}^{-1} + \bar{\mathbf{P}} + \mathbf{M}'_{ij}{}^{-1})^{-1} (\bar{\mathbf{B}}^{-1} - \frac{1}{2}\bar{\mathbf{P}}) \right] \mathbf{p}_{12} \right\} \tag{7}$$

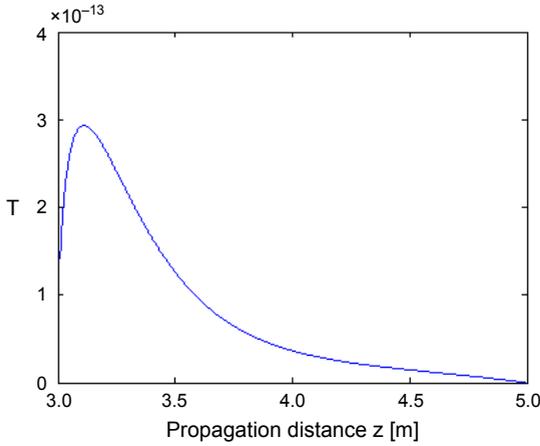


Fig. 1. The variation of  $T$  with power law  $\alpha$ , other parameters are  $L_0 = 10$  m,  $l_0 = 1$  mm,  $C_n^2 = 10^{-13} \text{ m}^{\alpha-3}$ .

where

$$\bar{\mathbf{B}} = \begin{bmatrix} z\mathbf{I} & 0 \\ 0 & -z\mathbf{I} \end{bmatrix}$$

$$\bar{\mathbf{P}} = \frac{2\pi^2 k z T}{3i} \begin{bmatrix} \mathbf{I} & -\mathbf{I} \\ -\mathbf{I} & \mathbf{I} \end{bmatrix}$$

and  $\boldsymbol{\rho}_{12}^T = (\boldsymbol{\rho}_1^T, \boldsymbol{\rho}_2^T) = (x_1, y_1, x_2, y_2)$ ,  $\bar{\mathbf{I}}$  is  $4 \times 4$  unitary matrix, and  $\mathbf{I}$  is  $2 \times 2$  unitary matrix; Det is the determinant of a matrix.

It can be deduced from Eq. (7) that the anisotropic electromagnetic beam in the output plane in turbulence media depends on the parameters of the source, the perturbation of turbulence, and the propagation distance.

### 3. Generalized Stokes parameters and polarization properties

The polarization properties of an electromagnetic beam at a point in space can be determined by using Stokes parameters which have been generalized from one-point quantities to two-point counterparts recently [20]. They contained not only polarization properties but also coherence properties of beams [21] and can be measured by Young's interference experiment [22, 23]. The changes in generalized Stokes parameters in the optical system or random media are also an interesting problem which has attracted a lot of attention [24, 25]. The four parameters can be expressed as follows:

$$S_i(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2) = \text{Tr} \left[ \mathbf{W}(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2, \boldsymbol{\omega}) \cdot \boldsymbol{\sigma}_i \right], \quad i = 0, 1, 2, 3 \quad (8)$$

where  $\mathbf{W}$  is the Pauli spin matrices,  $\sigma_0$  is the unit matrix:

$$\sigma_0 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \sigma_1 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \sigma_2 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \sigma_3 = \begin{pmatrix} 0 & -i \\ -i & 0 \end{pmatrix}$$

Substituting Eq. (7) into Eq. (8), we can get the changes in four generalized Stokes parameters of anisotropic beams propagation through turbulence media. Figure 2 illustrates the changes in normalized Stokes parameters of two pairs of points on the output plane of anisotropic beams traveling through turbulence with different power law  $\alpha$ , other turbulence parameters are the same as those which we used in Fig. 1. The parameters of the source which we used in this paper are:  $A_x = 3$ ,  $A_y = 1$ ,  $B_{xx} = B_{yy} = 1$ ,  $B_{xy} = 0.2\exp(i\pi/6)$ ,  $B_{yx} = 0.2\exp(-i\pi/6)$ ,  $\delta_{xx} = \delta_{yy} = 1.5$  mm,  $\delta_{xy} = \delta_{yx} = 2$  mm, and  $\lambda = 632.8$  nm.

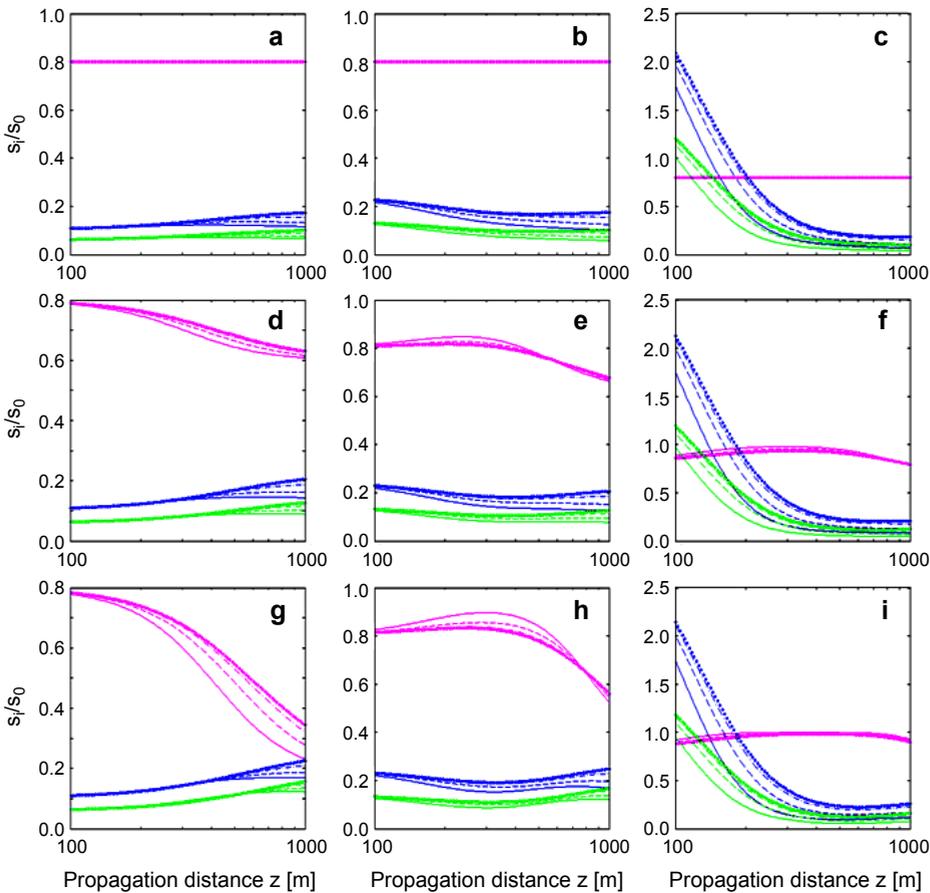


Fig. 2. Changes in normalized generalized Stokes parameters with the propagation distance of an anisotropic electromagnetic beam in turbulence. Two spatial points in the output plane are:  $\rho_1 = (0, 0)$ ,  $\rho_2 = (0, 0)$  (a, d, g);  $\rho_1 = (0.002, 0)$ ,  $\rho_2 = (0, 0)$  (b, e, h);  $\rho_1 = (0.003, 0)$ ,  $\rho_2 = (0, 0)$  (c, f, i).

In Fig. 2, the changes in normalized Stokes parameters in modified non-Kolmogorov turbulence have been illustrated. The pink line expressed the changes in  $s_1$ . The blue line expressed the changes in  $s_2$ , the green line expressed the changes in  $s_3$ . Different line types show the change in the parameters under different power laws of turbulence: the solid line shows the changes in generalized Stokes parameters under the power law of  $\alpha = 3.11$ , the dashed line is under the power law of  $\alpha = 3.4$ , the dash-dotted line is under the power law of  $\alpha = 3.7$ , the dotted line shows the changes in Stokes parameters in free space ( $T = 0$ ). In Figures 2a–2c  $\sigma_x$  and  $\sigma_y$  are assumed to be equal  $\sigma_x = \sigma_y = 0.2$  mm, and we find that as the propagation distance increases,  $s_1$  is stable for different pairs of spatial points on the output plane, and  $s_3$  will fluctuate due to turbulence perturbation and these fluctuations will be different from each other under different turbulence power laws. These variations are also different from each pair of points. In Figures 2d–2i we show the changes in the parameters of the beams generated by anisotropic sources, *i.e.*, the sources with  $\sigma_x = 0.2$  mm,  $\sigma_y = 0.3$  mm for Figs. 2d–2f, and  $\sigma_x = 0.2$  mm,  $\sigma_y = 0.5$  mm for Figs. 2g–2i.

In these situations,  $s_1$  shows more fluctuation features as the deviation of spot beam width increases ( $\Delta = \sigma_x - \sigma_y$ ). The variations of other parameters are less influenced by the deviation  $\Delta$ . However, the turbulence perturbation affects the changes in  $s_2$  and  $s_3$  much more than the changes in  $s_1$  in Figs. 2a–2f. In Figures 2g and 2h, the intensity of turbulence influences the change in  $s_1$  obviously, but in Fig. 2i, the changes in  $s_1$  are less influenced by the turbulence. From numerical simulation results, we can summarize the characteristics of generalized Stokes parameters of stochastic beams: for isotropic beams,  $s_1$  will hold a constant during the propagation whether in turbulence or free space,  $s_2$  and  $s_3$  will change in different turbulence conditions through the propagation. The trends of these changes depend on the positions of the two points in the output plane. For anisotropic beams,  $s_1$  will fluctuate during the propagation; these fluctuations are influenced by the turbulence. We noticed that in Fig. 2d–2f, the changes in  $s_2$  and  $s_3$  are similar to the changes in the same parameters in Figs. 2a–2c, but in Figs. 2g–2i, the changes in these two parameters are different. It means that the changes in  $s_2$  and  $s_3$  will be similar to each other if the deviation  $\Delta$  is small, but there is a significant difference between them when  $\Delta$  increases to a certain extent.

From the discussion by VOLKOV *et al.* [2], the relationship of generalized Stokes parameters and SDCP is:

$$P(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2) = \sqrt{s_1^2(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2) + s_2^2(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2) + s_3^2(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2)} \quad (9)$$

If  $\boldsymbol{\rho}_1 = \boldsymbol{\rho}_2 = \boldsymbol{\rho}$ , SDCP is equal to the degree of polarization (DOP) of the beam. In Fig. 2a, we can notice that  $s_2$  and  $s_3$  will reach the initial value if the propagation distance is long enough,  $s_1$  will be a constant, so it can be deduced that DOP of the isotropic beam will return to its initial value when the beam propagates sufficiently long distance; the phenomenon is the so-called polarization self-reconstruction. But for an anisotropic beam, or for two-points SDCP cases like those presented in other

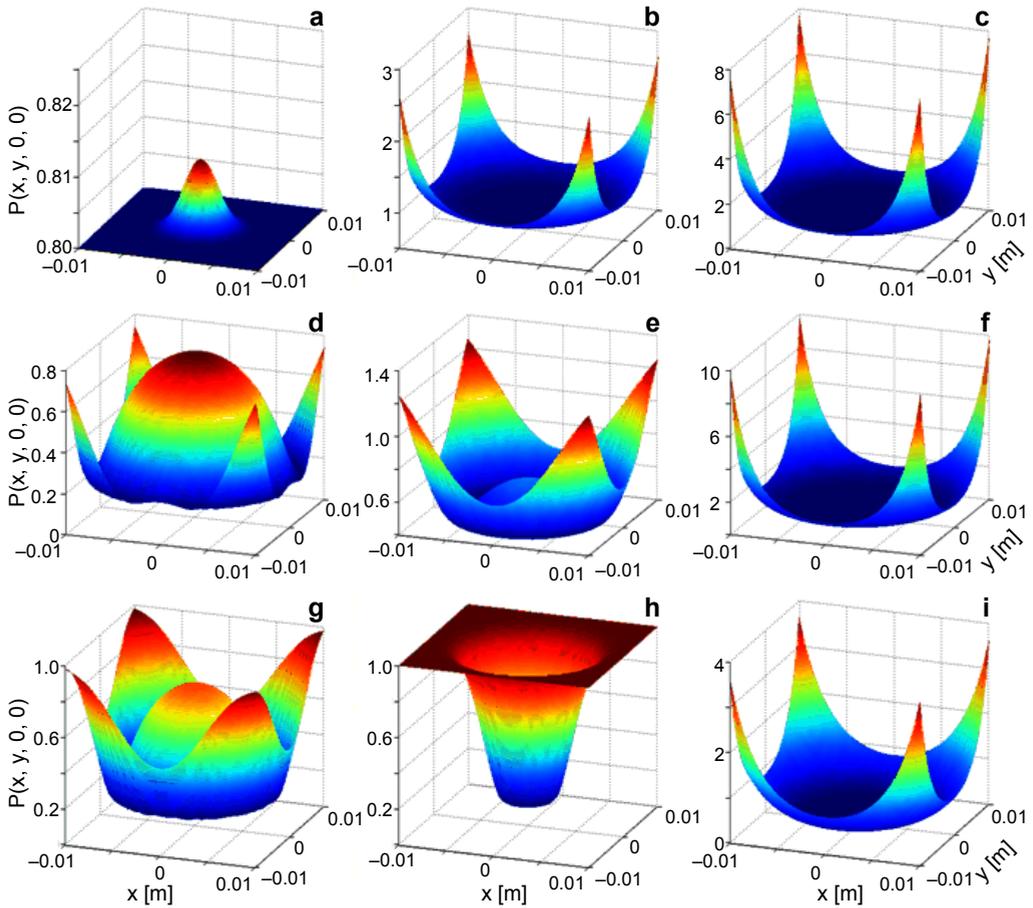


Fig. 3. Profiles of the degree of cross-polarization  $P(x, y, 0, 0)$  of an anisotropic electromagnetic beam in modified turbulence  $\alpha = 3.11$ ;  $z = 10^2$  m (**a**, **d**, **g**),  $z = 10^3$  m (**b**, **e**, **h**), and  $z = 3 \times 10^3$  m (**c**, **f**, **i**).

figures in Fig. 2a, there is no such property. In order to study the evolution of transverse SDCP through turbulence, Fig. 3 is organized to show the two-dimensional distribution of SDCP of stochastic beams with the same source parameters setting as in Fig. 2. We can see that the central values of SDCP of the isotropic beam are much greater than circumference values at the propagation distance of  $z = 0.1$  km, but for the anisotropic beam, the center and four corners ( $\pm 0.01, \pm 0.01, 0, 0$ ) of the values are greater than the remaining value of the area.

If the propagation distance is long enough, the profile of the distribution of transverse SDCP of the isotropic beam and the anisotropic beam will be the same. The SDCP values in the surrounding region will increase as the propagation distance increases, the final SDCP distribution profile is almost the same for the isotropic beam

and the anisotropic beam, they only differ in quantity. The four corners of the SDPC will get to the maximum value at  $z = 3$  km. The power law of turbulence is assumed as  $\alpha = 3.11$ , it means the perturbation is relatively strong in this situation.

In order to obtain how the turbulence perturbation affects the changes in SDPC, we show the variation of SDPC at different propagation distances with the difference power law of modified non-Kolmogorov turbulence in Fig. 4. These figures illustrate that the influence of atmospheric turbulence on the distribution of SDPC becomes obvious as the propagation distance increases. From Figs. 4a–4c to Figs. 4g–4i, we can see that the profile of transverse SDPC fluctuates much more in strong turbulence

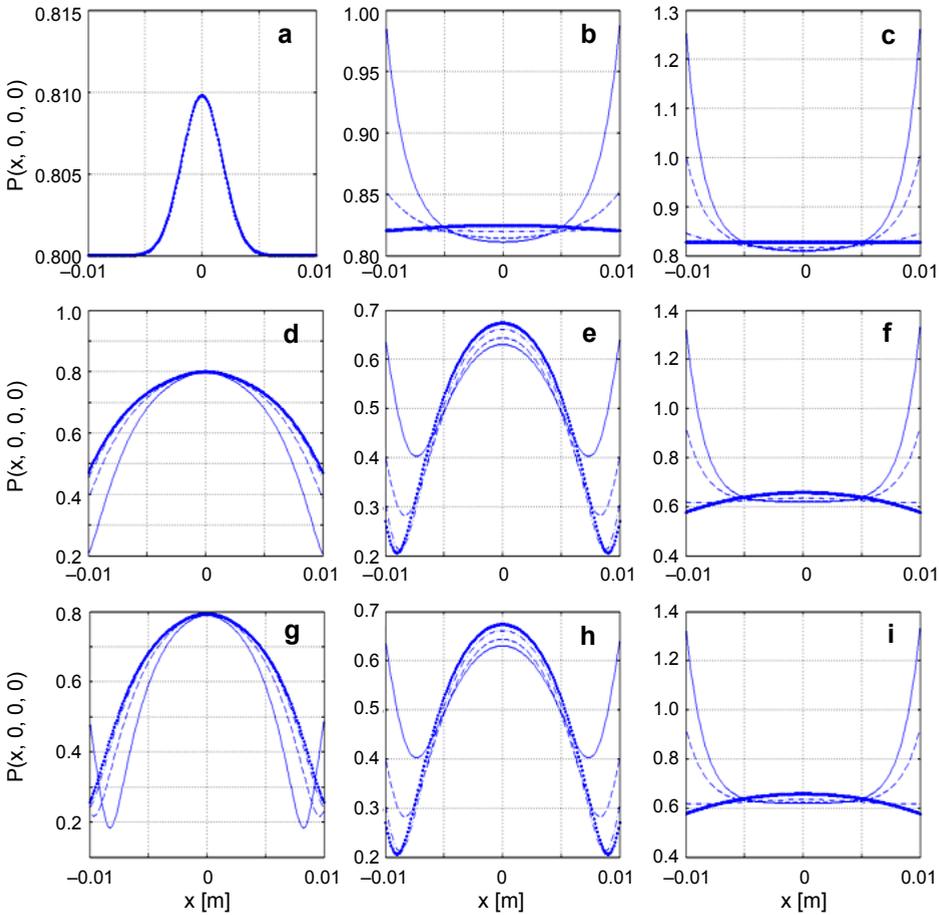


Fig. 4. Profiles of one-dimensional degree of cross-polarization  $P(x, 0, 0)$  of an anisotropic electromagnetic beam in modified turbulence with different propagation distance and power law of turbulence.  $z = 10^2$  m (a, d, g),  $z = 10^3$  m (b, e, h), and  $z = 3 \times 10^3$  m (c, f, i). Solid line –  $\alpha = 3.11$ , dashed line –  $\alpha = 3.4$ , dash-dotted line –  $\alpha = 3.7$ , dotted line – free space.

than that in weak turbulence and free space during the propagation. The final profiles of SDCP of the isotropic and anisotropic beam in weaker turbulence also show more flat features than those in stronger turbulence.

## 4. Conclusions

In summary, the evolution of generalized Stokes parameters of stochastic beams propagation in modified non-Kolmogorov atmospheric turbulence was investigated, and the changes in SDCP of the beams were carried out directly on the basis of these parameters. The source parameters were chosen as an isotropic case and anisotropic case. Our results indicated that the profiles of distribution of SDCP of these two kinds of stochastic beams are different in the near field, as the propagation distance increases. The influence of turbulence becomes obvious on the variations of the distribution, at last, the difference in the profiles of the distribution of SDCP between the isotropic beam and the anisotropic beam will disappear if the propagation distance is long enough. The difference in the sources only determine the quantity of final SDCP of the beam but the profiles of the distribution of SDCP mainly determine the turbulence perturbation. The research may be helpful for the application in optical communication, in imaging or remote sensing systems in the atmosphere.

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